# The structure of ELMS and the distribution of transient power loads in MAST



## **Presented by Andrew Kirk EURATOM / UKAEA Fusion Association**

- The filamentary structure of ELMs
- The effect of rotation
- The radial expansion of ELMs
- Filaments during disruptions



#### **First evidence for radial efflux**

TS profiles obtained near ELM peak show the formation of a broad  $n_e$  tail only on the outboard side



ELMs show  $D_{\alpha}$  emission well beyond outboard separatrix





Large particle flux (J<sub>SAT</sub>) to outboard midplane probe during ELM





#### **First evidence for ELM structure**



 $V_{\phi}^{ped} \sim 15 \text{ kms}^{-1}$ 



When the Reciprocating Probe is close to the plasma edge - structure is observed in the  $J_{sat}$  signal



#### The effect of toroidal rotation





$$V_{\phi}^{ped}$$
~25 kms<sup>-1</sup>



#### The effect of rotation - toroidal mode number





Range of ELMs give mean mode number ~ 10

Consistent with most unstable modes typically predicted by linear theory (ELITE)





## **ELM spatial structure** (experiment)





Filaments are observed which are consistent with the structures expected from the theory of the non-linear evolution of ballooning modes [Wilson and Cowley]



### ELM spatial structure (experiment+theory)



Image simulation of the expected structure with  $q_{95}=4$  and n=10







### The spatial and temporal evolution of an ELM



ΨN



 $\bigcirc$ Radius Radius Radius



arb. units

MAST

Time (s)

Theory

## The spatial and temporal evolution of an ELM



- simulated







#### The effect of rotation - edge velocity shear

Impurity velocity using visible Doppler spectroscopy

125  $\mu$ s time resolution



Edge velocity shear before ELM

MAST

shear greatly reduced at ELM peak

then quickly recovers



#### **The effect of rotation - edge velocity shear**

MAST



Edge velocity shear greatly reduced at ELM peak

A necessary requirement if the filaments are to propagate radially



#### The radial expansion of the filament







- The particle flux (Jsat) delayed from divertor  $D_{\alpha}$  increases with distance outside the separatrix  $(\Delta_r)$
- Estimated radial velocity  $V_r \sim 1 \text{ km/s}$



#### The radial expansion of the filament





The time delay measured has two components the true radial expansion velocity and the delay due to the finite toroidal extent of the localised filaments

#### The true radial expansion velocity can be estimated using a simulation



#### **Delay depends on radial and toroidal velocity**







#### The radial expansion of the filament



Monte Carlo simulation of calculated  $V_r$  as a function of  $\Delta_r$  for:



#### The radial expansion of the filament





1.4

1.2

1.0

0.8

0.6

0.4

(kms<sup>-1</sup>)

>



MAST data

Accelerating filaments are expected from the non-linear theory [Wilson and Cowley]







Using two toroidally separated

probes located at the same

distance from the plasma edge

#### Plan view of MAST





in Europe

shot 10658





shot 10658







Results from simulation

using the same

parameters as used for

the radial velocity

Giving further confidence in the n~ 10 filament description of ELMs



#### **Filaments as a pre-cursor to disruptions**



The non-linear ballooning mode theory may also play a role in some types of disruptions:

 $\Rightarrow$  A low n mode grows near the plasma edge

 $\Rightarrow$  exceeding the ballooning mode threshold locally

 $\Rightarrow$  produces a radially propagating flux tube connected to the core

 $\Rightarrow$  on contact with plasma facing components - releases impurities that trigger a thermal quench



#### **Filaments as a pre-cursor to disruptions**



#### Possible evidence for these filaments



Visible images of single (n=1) filaments observed prior to some disruptions



# IR images of the interaction of filaments with vessel components









# IR images of the interaction of filaments with vessel components





Start of thermal quench

15 ms later



#### **Summary**



• There is evidence that filament-like structures exist during ELMs and some types of disruptions, consistent with the predictions of the non-linear evolution of ballooning theory

• In ELMs the filaments are generated on a 100  $\mu$ s timescale, rotate with the outboard edge of the plasma and accelerate out from the outboard side.

• The importance of these filaments is that they can provide toroidally and poloidally localised heat fluxes to plasma facing components

