

# **Neon transport in JET plasmas close to ITB formation with monotonic and reversed q-profiles.**

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\* See appendix of J. Pamela, Proc. IAEA Conf. (Sorrento 2000)

## **1. Introduction**

Recently, optimised shear research at JET has concentrated on two different types of scenario: one with LHCD preheat resulting in a negative magnetic shear and the other with an ohmic preheat leading to a weak positive magnetic shear. The ITB formation and characteristics of these two types of discharge have been studied by systematic variation of input power and torque [1]. In this paper, we study the impact on impurity transport of scenarios for weak positive and negative magnetic shear ITB plasmas.

## **2. Experimental technique**

In this series of discharges, small quantities of neon (less than 0.1%) have been injected during the whole duration of the main heating phase. Ne was injected by edge gas fuelling located on the machine midplane. The valve opening had to be modulated for the gas fuelling system to deliver such a small amount on average. Neon edge lines were not detected by VUV spectroscopy. Neon densities have been measured by charge-exchange recombination spectroscopy (NeX, 5248.5 Å).

Impurity transport coefficients were deduced from a least-squares fit of neon densities modelled by a  $1^{1/2}$ D transport model code (SANCO) to the measured densities. The neon influx to the plasma is not measured but can be simulated from the pressure drop in the reservoir and knowledge of the fuelling system geometry. The Ne gas is assumed to diffuse with a flux proportional to  $\text{grad}(n^2)$ ,  $n$  being the density of gas in the fuelling system. Another issue in the determination of transport coefficients is the least-squares fit program. The latter limits the definition of D and V/D profiles to a (changeable) grid of only 4 radial positions connected by straight lines. The least-squares fit program has a constant D and V/D for defined time windows. A smooth evolution of D and V/D from a time window to the next one is not possible at present.

#51573

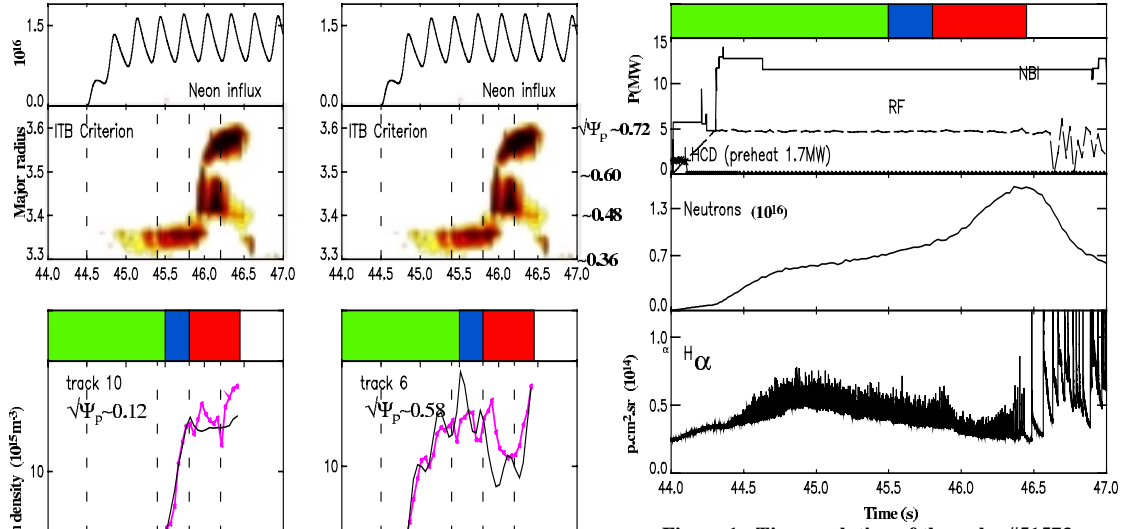


Figure 1: Time evolution of the pulse #51573

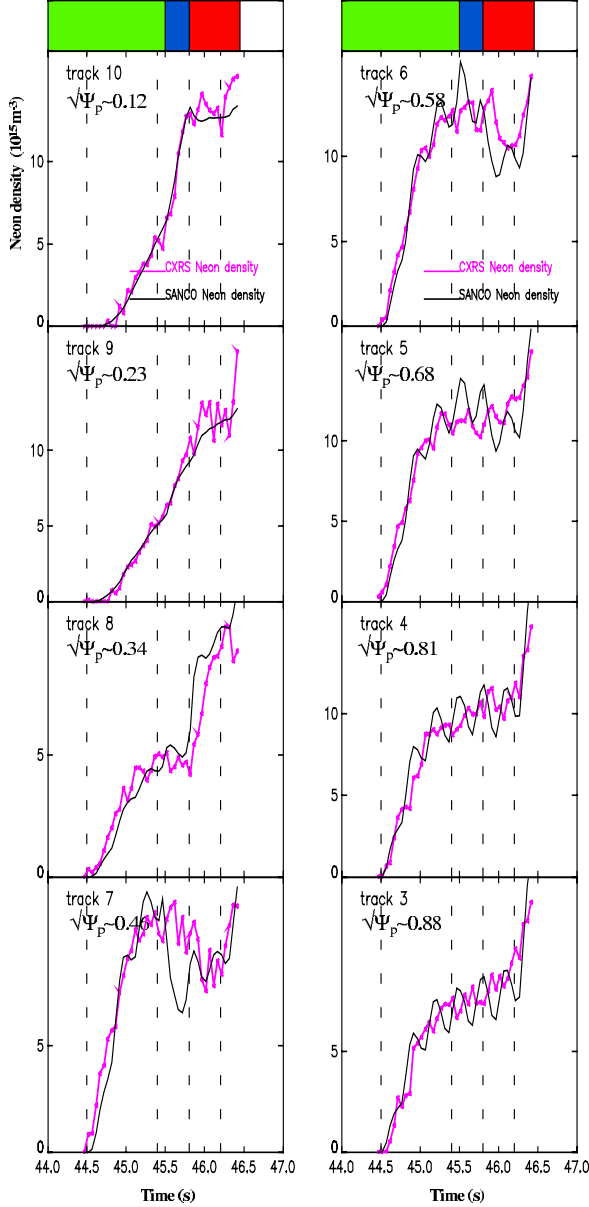


Figure 2: (from top to bottom) **Simulated neon gas influx** to the plasma used in SANCO. **ITB Criterion:** contour plots of  $\rho_T^*$ , [2], showing the strength and position of the barrier: the darker the colour, the higher  $\rho_T^*$ , the stronger the ITB. **Defined time slices for  $\chi^2$  fit:** the coloured rectangles indicate the time slices used in the  $\chi^2$  fit. **Time evolution of NeX density** measured by CXRS in the different CCD tracks (in pink) compared to SANCO modelled density (in black).  $\Psi_p$  is the normalised poloidal flux

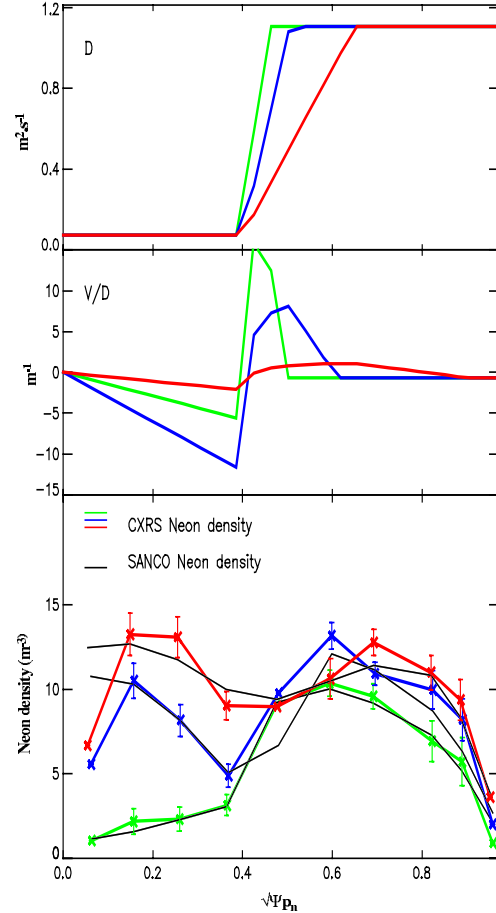
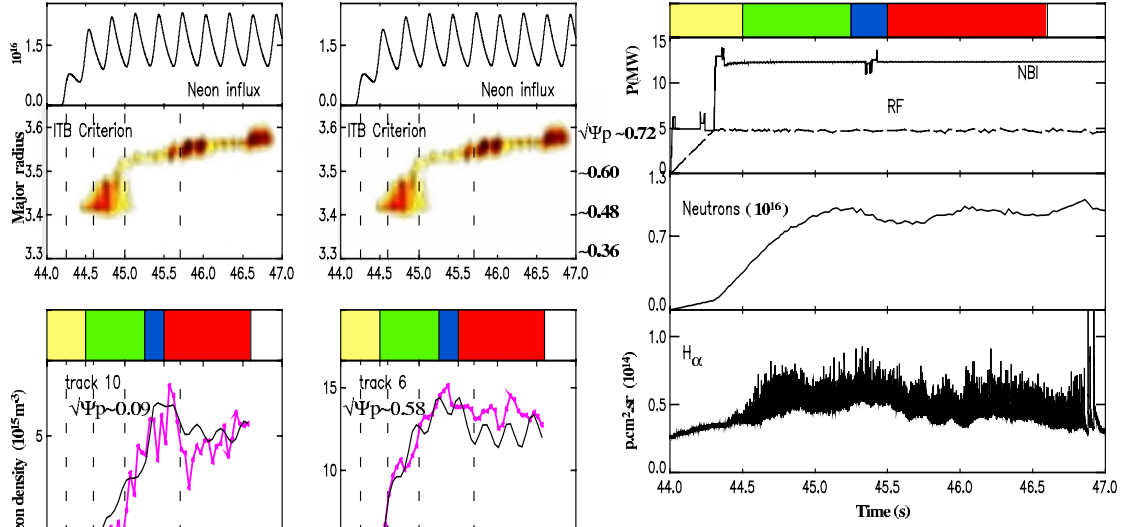
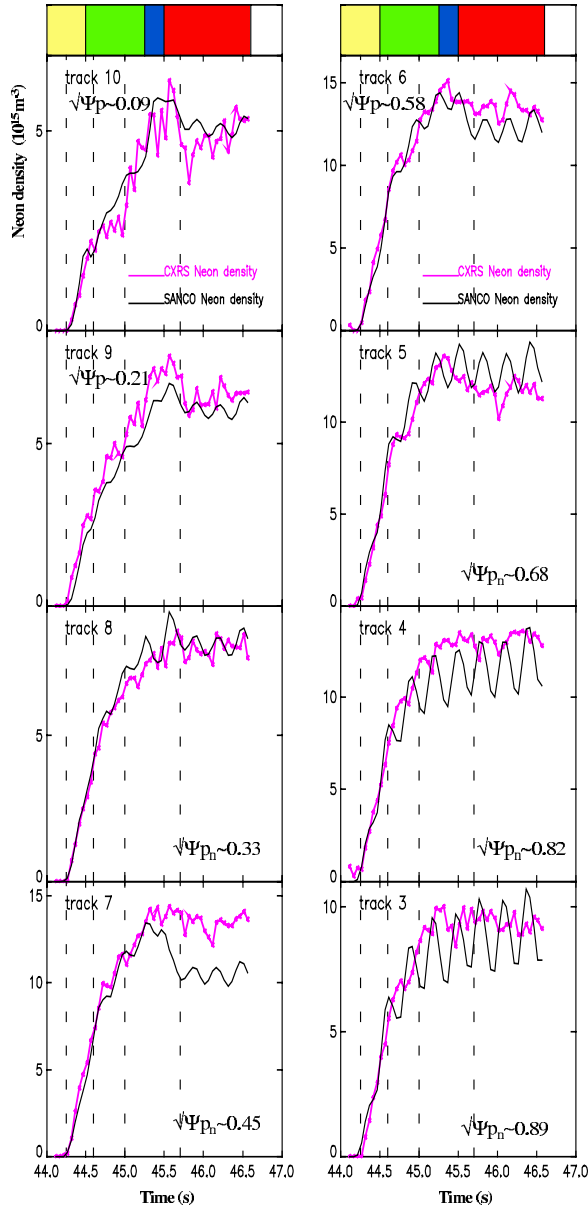


Figure 3: (from top to bottom) **Profiles of the diffusion coefficient (D) and V/D against  $\sqrt{\Psi_{pn}}$**  used in the  $\chi^2$  fit, V being the convection coefficient and  $\Psi_{pn}$  the normalised poloidal flux. **Typical measured and modelled NeX densities** against  $\sqrt{\Psi_{pn}}$  during the different time slices. Tracks plotted without error bars were not considered in the  $\chi^2$  fit.

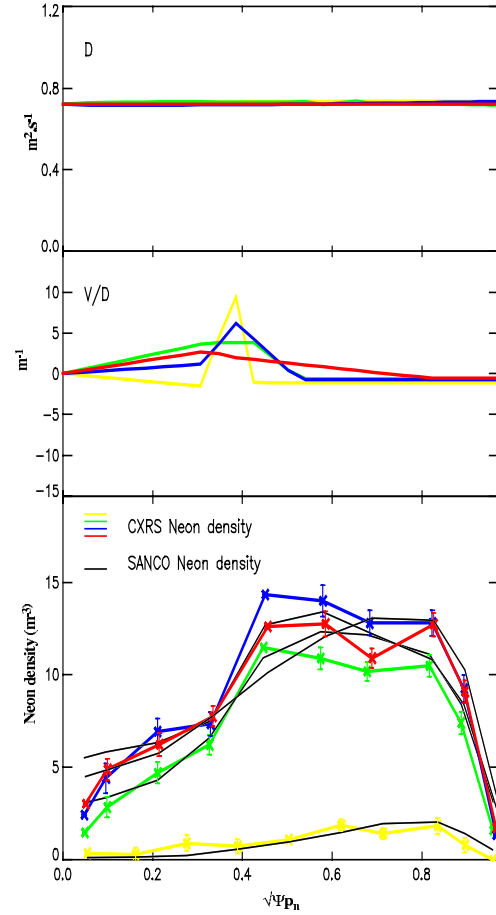
# #51598



**Figure 4: Time evolution of the pulse #51598**



**Figure 5: As in figure 2 for #51598**



**Figure 6: As in figure 3 for #51598**

### 3. Determination of impurity transport coefficients

The transport coefficients are determined for pulses #51573 and #51598. These discharges correspond to a negative and weak positive magnetic shear ITB plasma, respectively, close to ITB formation at 17MW input power with a dominant normal beam heating. Figures 2, 3, 5 and 6 show the comparison of the experimental data with the SANCO modelled densities.

**Justification for the fit #51573:** The delayed appearance of neon in tracks 8, 9 and 10 compared to the rest of the tracks shows a strongly reduced diffusion coefficient in the plasma core.  $D$  is determined at  $\sim 44.5$ s and kept fixed for the rest of the discharge. A positive barrier for  $V/D$  is required at  $\sqrt{\Psi}p_n \sim 0.45$  to reproduce the steep gradient between tracks 6,7 and 8, and a negative convection is needed for  $\sqrt{\Psi}p_n \leq 0.45$  to explain the time dependence of the density in tracks 8,9 and 10. A close look at the time dependence of Ne density in each track gives information of the ITB time behaviour. The sequence of events, seen in figure 2, can be explained by an expanding positive (outward convection)  $V/D$  region, as shown in figure 3.

**Justification for the fit #51598:** No delay in the appearance of Ne between tracks is observed. The diffusion is therefore kept constant in space, determined at  $\sim 44.2$ s and kept fixed in time. Here again, a positive barrier for  $V/D$  is required at  $\sqrt{\Psi}p_n \sim 0.4$  to reproduce the steep gradient between tracks 6,7 and 8, but for this pulse a positive convection is required at  $\sqrt{\Psi}p_n \leq 0.4$  to reproduce the density time dependence in tracks 8,9 and 10. Here again, an expanding  $V/D$  positive (outward convection) region can explain the density time evolution in each track.

### 4. Conclusion

We show clear evidence of the dependence on impurity transport of the chosen magnetic shear scenario. Neon accumulates in reversed magnetic shear ITB plasmas, due to inward convection, but not in weak positive shear ITB plasmas, where convection is outwards. Similar conclusions were drawn from the study of the equivalent discharges at 15 MW input power (#51594, #51595). This result is in agreement with the impurity behaviour in high performance discharges [3]. The implications of this work are two-fold. Firstly, there is little hope in controlling the impurity content in negative magnetic shear ITB plasmas. Secondly, there will be no impurity accumulation in weak positive magnetic shear ITB plasmas if the outward convection can be maintained on a longer time scale.

### 5. Reference

- [1] C.Challis, et al., at this conference.
- [2] G. Tresset, et al, at this conference.

- [3] R. Dux, et al, at this conference